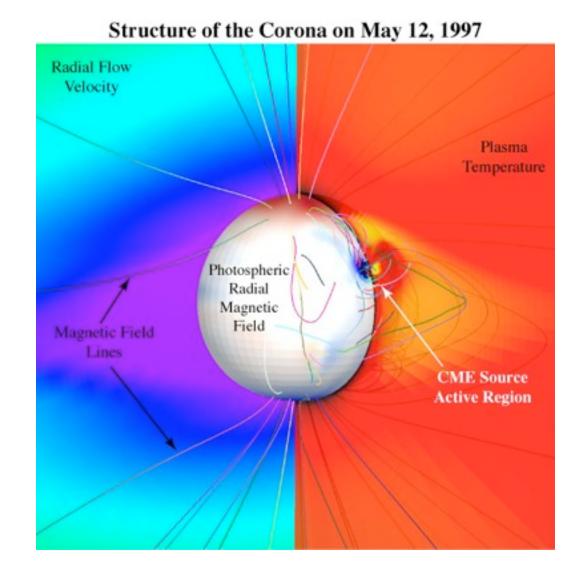
Using LOS Magnetograms in Global Heliospheric MHD Models

Pete Riley, Zoran Mikic, Jon A. Linker, Roberto Lionello

Predictive Science, San Diego, CA.

Magnetogram Workshop UCLA April 2-3, 2009





Some Acronyms:

- MAS: Magnetohydrodynamics Around a Sphere (The PSI coronal MHD code)
- Enlil: God of Wind (NOAA/SEC Heliospheric code)
- CORHEL: CORona and HELiosphere (Coupled MAS, WSA, PFSS, PFCS, Enlil, MASIP Package)
- MASIP: MAS in the InterPlanetary medium
- WSA, PFSS, PFCS: More acronyms. Will it ever end?

A better acronym of "ADAPT"

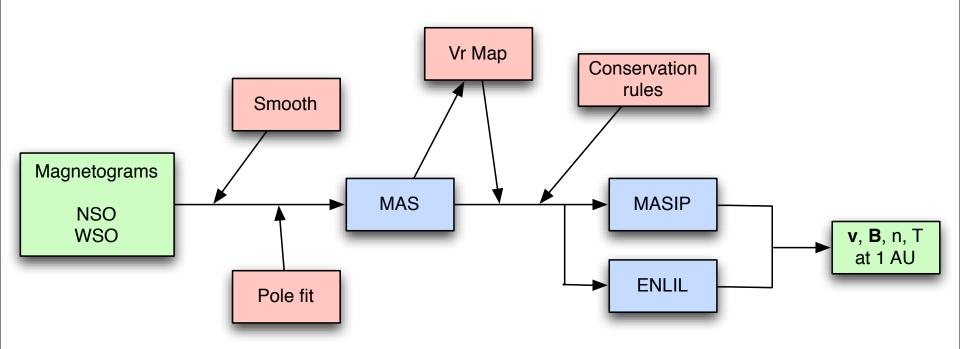
 ADAPT: Air Force data Assimilation for Photospheric Flux Transport

 AFART: Air Force Assimilation of Flux and Transport

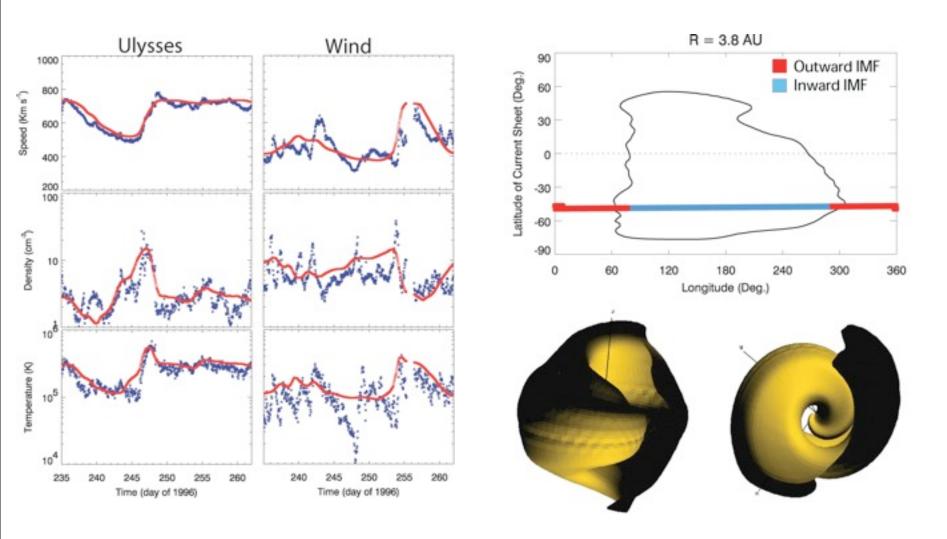
The 'old' version of CORHEL (MAS really) relies on a polytropic relationship between pressure and density

$$\begin{array}{rcl} \nabla\times\mathbf{A} &=& \mathbf{B},\\ \frac{\partial\mathbf{A}}{\partial t} &=& \mathbf{v}\times\mathbf{B} - \frac{c^2\eta}{4\pi}\nabla\times\mathbf{B},\\ \frac{\partial\rho}{\partial t} + \nabla\cdot(\rho\mathbf{v}) &=& 0,\\ \frac{1}{\gamma-1}\left(\frac{\partial T}{\partial t} + \mathbf{v}\cdot\nabla T\right) &=& -T\nabla\cdot\mathbf{v},\\ \rho\left(\frac{\partial\mathbf{v}}{\partial t} + \mathbf{v}\cdot\nabla\mathbf{v}\right) &=& \frac{\nabla\times\mathbf{B}\times\mathbf{B}}{4\pi} - \nabla p + \rho\mathbf{g} + \nabla\cdot(\nu\rho\nabla\mathbf{v}),\\ \gamma &=& 1.05. \end{array}$$

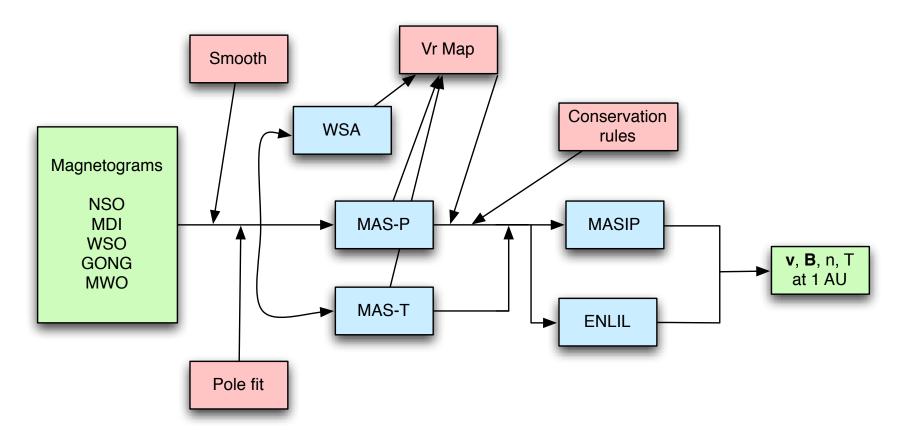
CORHEL 3.5/4.0



Comparisons with in situ data demonstrate that that the model can reproduce essential observations



Improvements to CORHEL currently under Development





The latest version of MAS includes energy transport processes

$$\nabla \times \mathbf{B} = \frac{4\pi}{c} \mathbf{J}$$

$$\nabla \times \mathbf{E} = -\frac{1}{c} \frac{\partial \mathbf{B}}{\partial t}$$

$$\mathbf{E} + \frac{1}{c} \mathbf{v} \times \mathbf{B} = \eta \mathbf{J}$$

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = 0$$

$$\rho \left(\frac{\partial \mathbf{v}}{\partial t} + \mathbf{v} \cdot \nabla \mathbf{v} \right) = \frac{1}{c} \mathbf{J} \times \mathbf{B} - \nabla p - \nabla p_w + \rho \mathbf{g} + \nabla \cdot (\nu \rho \nabla \mathbf{v})$$

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \mathbf{v}) = (\gamma - 1)(-\rho \nabla \cdot \mathbf{v} - \nabla \cdot \mathbf{q} - n_e n_p Q(T) + H)$$

$$\gamma = 5/3$$

$$\mathbf{q} = -\kappa_i \hat{\mathbf{b}} \hat{\mathbf{b}} \cdot \nabla T \qquad \text{(Close to the Sun, } r \leq 10R_s\text{)}$$

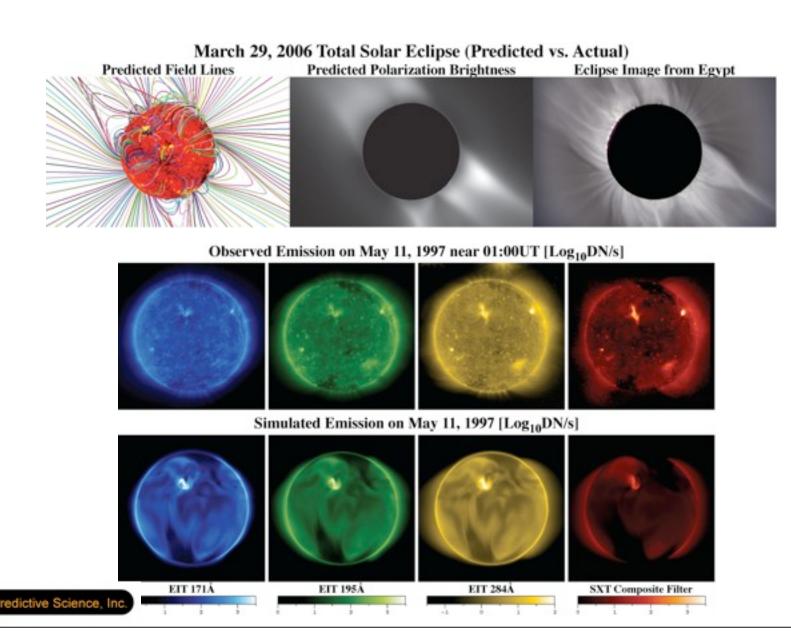
$$\mathbf{q} = 2\alpha n_e T \hat{\mathbf{b}} \hat{\mathbf{b}} \cdot \mathbf{v}/(\gamma - 1) \qquad \text{(Far from the Sun, } r \geq 10R_s\text{)}$$

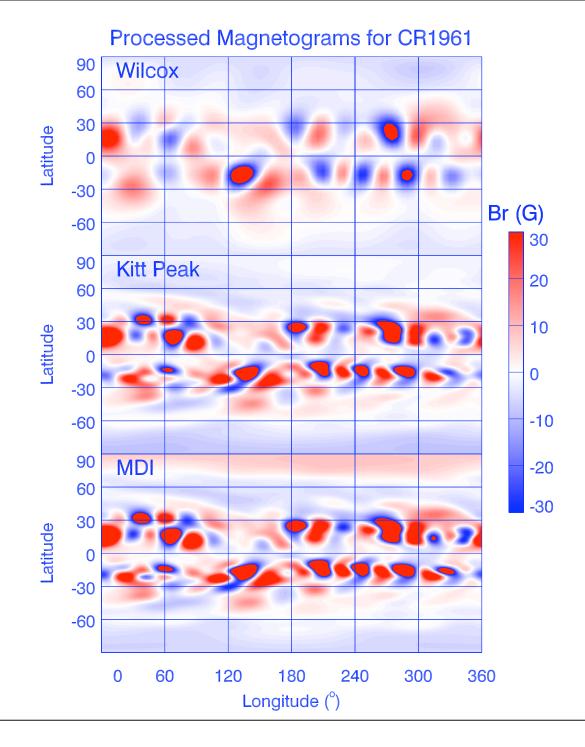
$$+ \text{WKB equations for Alfvén wave pressure } p_w \text{ evolution}$$

The new MAS is fully parallel with non-uniform grids in all three dimensions

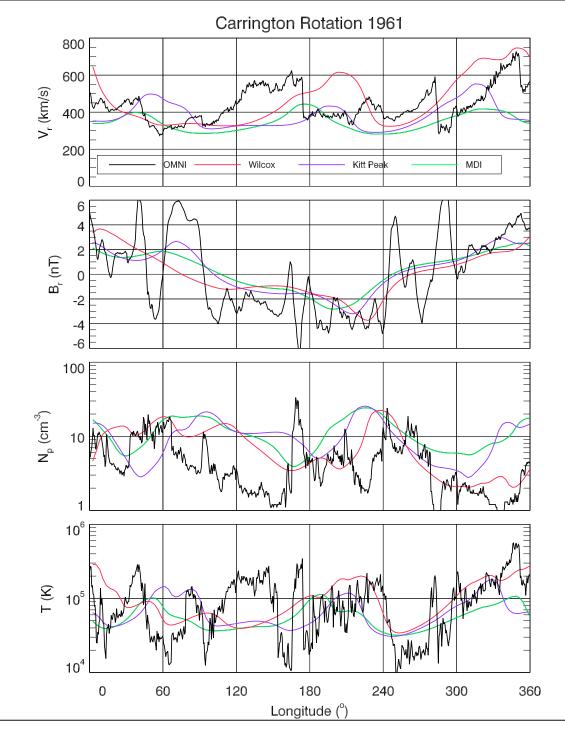
redictive Science, Inc.

The new MAS can reproduce white-light and emission measurements

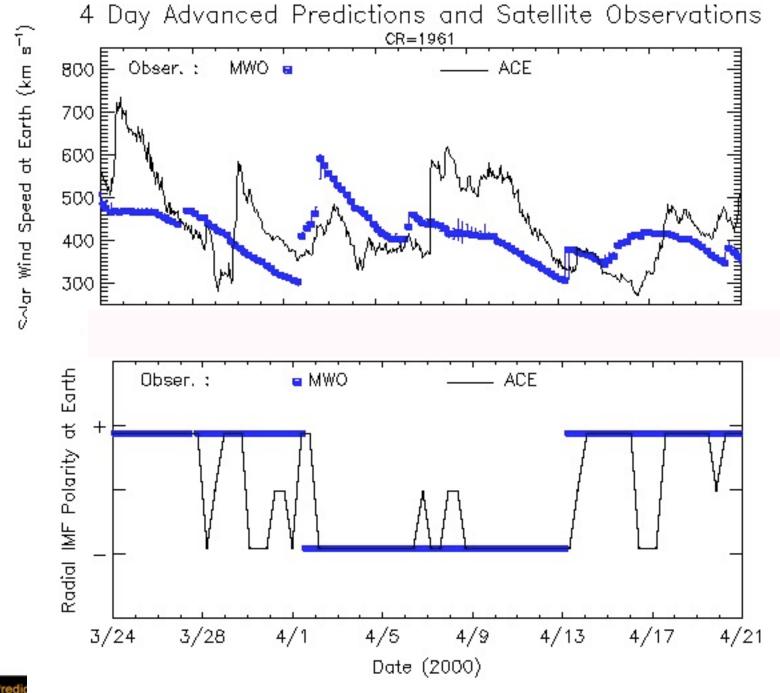




Predictive Science, Inc.



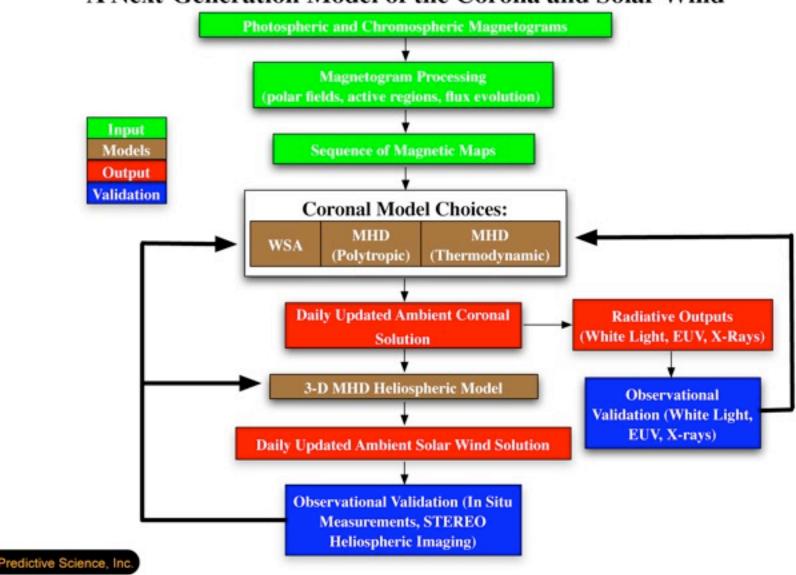
Predictive Science, Inc.





Our LWS Strategic Capability will eventually provide a near-real time SW model

A Next-Generation Model of the Corona and Solar Wind



In summary, MAS and CORHEL development is proceeding well. The quality, accuracy, and "ground truth" of the input magnetograms remains one of the most important unknowns.

Input magnetograms

- •"Ground truth" values
- Polar fitting
- •Far-side data
- Synchronic maps

Improved physical processes

- Thermal conduction
- Radiation
- Coronal heating
- Alfven Waves

Fully parallel code

- •Runs on Columbia, Ranger, & smaller clusters (e.g., Mac Pros)
- Higher resolution
- Non-uniform meshes

